

CPF calculation in circumlunar satellite laser ranging observations

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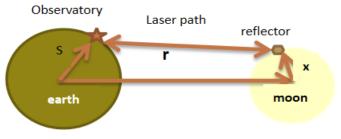




Context







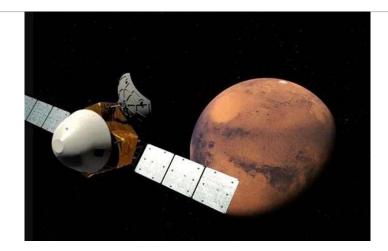
The mode of Lunar Laser Ranging



LLR at YNAO



The distribution of lunar retroreflectors



Launching artificial satellites to the vicinity of celestial bodies is a primary means for humans to explore target celestial bodies.

During the flight process, the artificial satellite needs to perform multiple orbital controls and maneuver to its predetermined orbit.

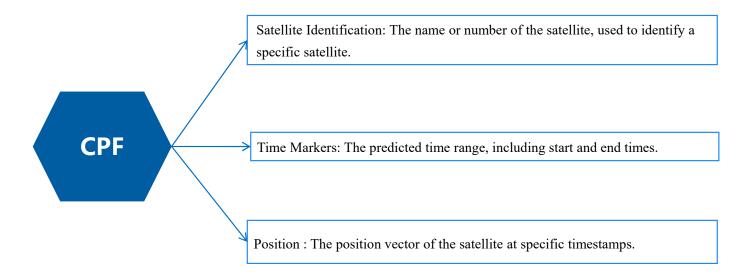
In this process, the determination of the satellite's orbit is crucial, and satellite laser ranging is the most effective technical means for precisely measuring the distance to the³ satellite.



Context

CPF (Consolidated Prediction Format) :

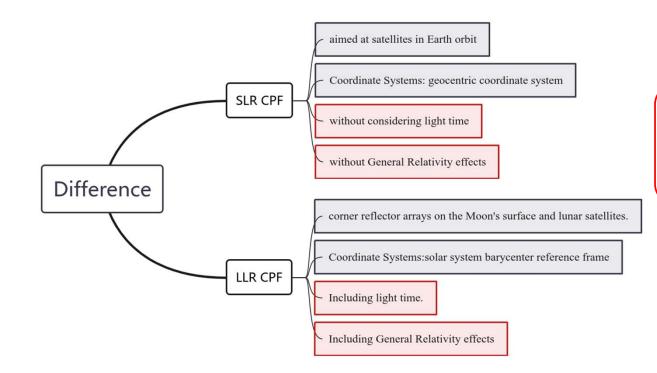
- CPF is a standardized file format used for storing and exchanging satellite orbit prediction data. The CPF file contains key information required for satellite laser ranging
- > The CPF file is a key component in satellite laser ranging activities, ensuring that ground stations can accurately point to the satellite and emit laser pulses at the right time for precise distance measurements.





Context

The differences between Lunar Laser Ranging (LLR) and Satellite Laser Ranging (SLR) in Consolidated Prediction Format (CPF) files.



For circumlunar satellites, the generation format of the LLR CPF files should be utilized.



Methods

Coordinate system transformation (ITRS to ICRS)

The coordinates of the station are in the ITRS, while the coordinates of the circumlunar satellite are described in the ICRS. Therefore, we need to uniform it to the same coordinate system.

Time conversion :(UTC to TDB)

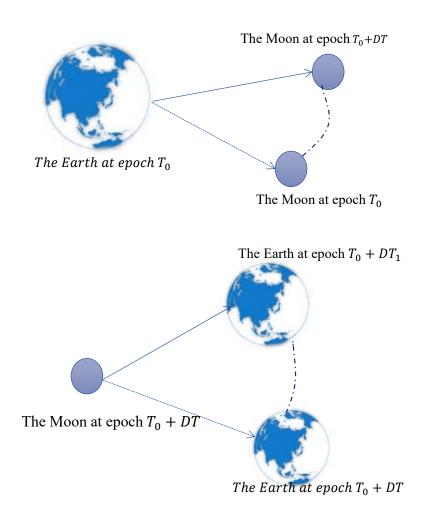
The observation station uses UTC, while the ephemeris is TDB time. It is necessary to unify the time to TDB.

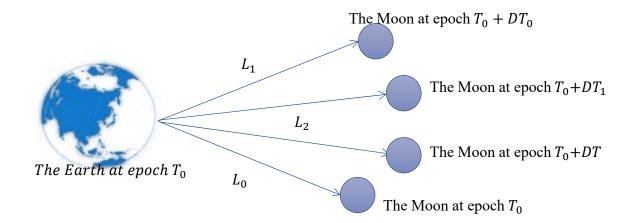
General Relativity effects:

Due to the gravitational effects of celestial bodies such as the Sun, Earth, and Moon, lasers will experience varying degrees of bending as they propagate through gravitational fields. Thereby increasing the light time. This experiment uses the gravitational bending equation from the IERS2010 Conventions to calculate this correction value. The formula is as follows:

$$\Delta t = \frac{\rho}{c} + \sum_{J} \frac{2GM_{J}}{c^{3}} \ln(\frac{r_{J1} + r_{J2} + \rho}{r_{J1} + r_{J2} - \rho})$$

Methods





(1) Starting from the geometric position at the given T_0 (initial) epoch, calculate the estimated light time DT_0 from the observer to the celestial body to obtain the initial distance L_0 .

(2) Based on the motion velocity of the celestial body, compute the new position at $T_0 + DT_0$ (Initial time plus light time) and recalculate the distance L_1 , estimating the light time DT_1 .

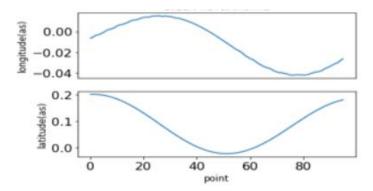
(3) Based on the motion velocity of the celestial body, compute the new position at $T_0 + DT_1$ and recalculate the distance L_2 , estimating the light time DT_2 .

(4) When DT_n - DT_{n-1} difference of light time is approximately equal to zero, the precise light time DT_n is determined, from which the accurate apparent position of the celestial body can be determined

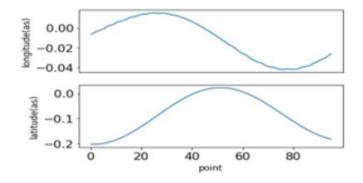




Conclusion



The forecasted launch position is differenced with the CPF file (with each time point spaced 15 minutes apart).



The forecasted reception position is differenced with the CPF file (with each time point spaced 15 minutes apart).

Conclusion: The difference between the laser transmission and reception positions and the CPF file published by the ILRS shows that the difference in the longitude direction in the geocentric coordinate system is 0.2 arcseconds, and the latitude difference is less than 0.1 arcseconds. Further experiments indicate that this difference does not increase when transformed into the observation station's coordinate system. Therefore, the differences in the longitude and latitude directions of the generated file are both better than the pointing accuracy of 1 arcsecond of the 1.2-meter telescope at the Yunnan Observatories, indicating the feasibility of the CPF file.



Conclusion

Generate CPF file of LRO(Lunar Reconnaissance Orbiter)				H1 CPF 2 YNO 2024 10 21 15 90 7 LRO imcce_inpop21 H2 103 103 0 2009 9 11 0 0 0 2009 9 12 0 0 0 900 0 1 0 0 0 3 H9		
				10 1 54708 0.0 0 -175172928.486 -323458590.004 96737370.32 10 2 54708 0.0 0 175234190.366 323497396.448 -96765008.38 30 1 -704. -35619. 13821. 25.6 10 1 54708 900.0 0 -195281050.717 -311694617.959 97112714.9 10 2 54708 900.0 0 195344800.880 311729463.931 -97140360.0 30 1 -3033. -35502. 13824. 25.6 10 1 54708 1800.0 0 -214602275.127 -298682305.078 97487490. 10 2 54708 1800.0 0 214668257.240 298713030.956 -97515142.2		
lrocpf_200909011_200909012	2024/10/21 23:27	文件	23 KB	30 1 -534835233. 13828. 25.6 10 1 54708 2700.0 0 -233059214.156 -284474102.415 97861694.		
Irocpf_200909012_200909013	2024/10/21 23:27	文件	23 KB	10 2 54708 2700.0 0 233127161.650 284500565.189 -97889353.0		
Irocpf 200909013 200909014	2024/10/21 23:27	文件	23 KB	30 1 -764234814. 13831. 25.6		
□ lrocpf_200909014_200909015	2024/10/21 23:27	文件	23 KB	10154708 3600.00 -250577965.421 -269127251.167 98235324. 10254708 3600.00 250647602.556 269149325.090 -98262990.0 301 -9904. -34245. 13835. 25.7 10154708 4500.00 -267088407.751 -252703551.720 98608379. 10254708 4500.00 267159450.749 252721128.903 -98636052.8 301 -12124. -33529. 13838. 25.7 -98636052.8 10154708 5400.00 -282524481.904 -235269114.511 98980857. 10254708 5400.00 282596640.098 235282105.431 -99008537.8 301 -14293. -32670. 13842. 25.7 10154708 6300.00 -296824454.983 -216894093.546 99352756. 10254708 6300.00 296897431.995 216902427.485 -99380443.3 301 -16401. -31670. 13845. 25.7 10154708 7200.00 310004662.319 1976552403.794 99724074. 10254708 7200.00 310004662.319 197656029.209 -99751768.0 <		

and can meet the needs for ranging of circumlunar satellites.

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Thank you teachers and professors for listening

